Video Denoising Based on a Spatiotemporal Gaussian Scale Mixture Model

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Abstract—We propose a video denoising algorithm based on a spatiotemporal Gaussian scale mixture (ST-GSM) model in the wavelet transform domain. This model simultaneously captures the local correlations between the wavelet coefficients of natural video sequences across both space and time. Such correlations are further strengthened with a motion compensation process, for which a Fourier domain noise-robust cross correlation algorithm is proposed for motion estimation. Bayesian least square estimation is used to recover the original video signal from the noisy observation. Experimental results show that the performance of the proposed approach is competitive when compared with state-of-the-art video denoising algorithms based on both peak signal-to-noise-ratio (PSNR) and structural similarity (SSIM) evaluations.

Index Terms—video denoising, image restoration, statistical image modeling, Gaussian scale mixture, cross correlation, motion estimation, Bayesian estimation

I. INTRODUCTION

Video signals are often contaminated by noise during acquisition and transmission. Removing/reducing noise in video signals (or video denoising) is highly desirable, as it can enhance perceived image quality, increase compression effectiveness, facilitate transmission bandwidth reduction, and improve the accuracy of the possible subsequent processes such as feature extraction, object detection, motion tracking and pattern classification.

Video denoising algorithms may be roughly classified based on two different criteria: whether they are implemented in the spatial domain or transform domain and whether motion information is directly incorporated. Spatial domain denoising is usually done with weighted averaging within local 2D or 3D windows, where the weights can be either fixed or adapted based on the local image content. A review of spatial domain filtering methods can be found in [1]. Transform domain methods first decorrelate the noisy signal using a linear transform (e.g., a wavelet transform), and then attempt to recover the transform coefficients of the original signal (e.g., by soft/hard thresholding [2] or Bayesian estimation [3]), followed by an inverse transform that brings the signal back to the spatial domain.

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The high degree of correlation between adjacent frames is a "blessing in disguise" for signal restoration. On the one hand, since additional information is available from nearby frames, a better estimate of the original signal is expected. On the other hand, the process is complicated by the presence of motion between frames. Motion estimation itself is a complex problem and it is further complicated by the presence of noise. In [1], performance of spatial domain motion compensated filters was evaluated. A multiresolution motion estimation scheme was proposed in [4] when the signal is corrupted with noise. In [5], a recursive filter is applied on wavelet transform coefficients along an estimated motion trajectory, where the filter taps are adaptively chosen based on the "reliability" of the motion vectors. Motion information or temporal correlations may also be incorporated by employing an advanced or adapted transform [6], [7] or by using an advanced statistical model that reflects the joint distributions of wavelet coefficients over space and time [8–10]. Some wavelet domain algorithms [5], [11] used robust motion indices that represent motion in an indirect way. Recently, a series of successful non-local patchbased methods emerged [12-17], where motion information is incorporated implicitly by adaptively clustering similar 2D or 3D patches. It was also shown that imposing sparseness prior models would further improve the performance of these algorithms [16], [17].

In recent years, there has been a growing interest in studying statistical models of natural images, which provide useful prior knowledge about natural images and play important roles in the design of Bayesian signal denoising algorithms [18]. While great effort has been made to study statistical models of static natural images [19], [20], much less has been done for natural video signals. In [21], the spatiotemporal Fourier power spectra of natural image sequences were investigated. In [22], independent component analysis was applied to local 3D blocks extracted from natural image sequences and the components optimized for independence are filters localized in space and time, spatially oriented, and directionally selective. Similar shapes of linear components were also obtained by optimizing sparseness via a matching pursuit algorithm [23]. In [24], it was observed that natural video sequences exhibit strong statistical prior of temporal motion smoothness, which can be captured by temporal local phase correlations in the complex wavelet transform domain. The application of statistical models for signal denoising has also been extended from 2D to 3D. In [8], a video denoising algorithm was proposed by extending the hiddern Markov tree model [25], [26] from 2D to 3D, which provides a simple but effective way to describe the inter-scale statistical correlations of wavelet coefficients.

In [9], a statistical model for wavelet coefficients of video signals was proposed, based on which maximum a posteriori (MAP) estimation and temporal filtering were used for video denoising.

In this paper, we propose a new video denoising method based on a spatiotemporal model of motion compensated wavelet coefficients. Our work was motivated by the success of the Gaussian scale mixture (GSM) model in static image denoising [18]. GSM was originally proposed in the statistics literature [27] and has been applied to the modeling of static natural images [28]. It was found to be a convenient and effective model to account for the non-Gaussian marginal distributions of wavelet coefficients, as well as the variancedependencies between neighboring wavelet coefficients. Since there exist strong correlations between wavelet coefficients across adjacent video frames, it is expected that grouping wavelet coefficients along temporal directions based on the GSM framework would help strengthen the statistical regularities of the coefficients and thus improve the performance of Bayesian signal denoising algorithms. It is also important to be aware that video signals are more than simple 3D extensions of 2D static image signals, where the major distinction is the capability of representing motion information [29]. If the motion between frames is properly compensated, then the temporal correlations between wavelet coefficients can be enhanced. Effective motion compensation relies on reliable motion estimation. In the literature, a large number of local motion estimation methods have been proposed, which are mostly based on block matching (e.g., [30]), optical flow [31], [32] or phase disparity [33] evaluations. To avoid local nonlinear motion compensation processes that affect the validity of the Gaussian noise model assumed in GSM denoising, we opt to use global motion estimation methods based on cross correlation [34-37]. The challenge here is that the motion information must be estimated from noisy video signals rather than the noise-free original signals. To address this issue, another important aspect of our work was the development and incorporation of a novel noise-robust motion estimation algorithm, which has not been carefully examined in previous video denoising algorithms.

II. BACKGROUND OF GSM IMAGE DENOISING

A random vector \mathbf{x} is a GSM if it can be expressed as the product of two independent components:

$$\mathbf{x} = \sqrt{z}\mathbf{u},\tag{1}$$

where \mathbf{u} is a zero-mean Gaussian vector with covariance matrix $\mathbf{C}_{\mathbf{u}}$, and z is called a mixing multiplier. The density of \mathbf{x} is then given by:

$$p_x(\mathbf{x}) = \int \frac{1}{[2\pi]^{N/2} |z\mathbf{C}_{\mathbf{u}}|^{1/2}} \exp(-\frac{\mathbf{x}^T (z\mathbf{C}_{\mathbf{u}})^{-1} \mathbf{x}}{2}) p_z(z) dz$$
(2)

where N is the size of the vector \mathbf{x} and $p_z(z)$ is the mixing density. When applying it for image modeling, \mathbf{x} is typically composed of a center wavelet coefficient, x_c , together with a set of coefficients located near x_c in the same wavelet subband or nearby subbands across scale and/or orientation. Among

various choices of the prior density distribution $p_z(z)$, the one that was found to give superior image denoising performance is the non-informative Jeffrey's prior [18], [38] given by

$$p_z(z) \propto \frac{1}{z}$$
 (3)

The key feature of this prior is its amplitude scale invariance, which means that the inference procedure is invariant with respect to changes in the measurement units (or the scale of amplitude) [38], [39].

Assume that the original image is contaminated by additive, independent white Gaussian noise, then in the wavelet transform domain, a noisy neighborhood coefficient vector \mathbf{y} can be modeled as

$$\mathbf{y} = \mathbf{x} + \mathbf{w} = \sqrt{z}\mathbf{u} + \mathbf{w}, \tag{4}$$

where \mathbf{w} is a noise coefficient vector with covariance matrix $\mathbf{C}_{\mathbf{w}}$. The group of neighboring coefficients constitutes a sliding window that moves across the wavelet subband. At each step, only the center coefficient, x_c , of the window is estimated (i.e., denoised). Consequently, the objective here is converted to estimating x_c of \mathbf{x} , given the noisy observation \mathbf{y} .

It is not difficult to show that the Bayes Least Square (BLS) estimator (which minimizes the expected value of the squared estimation error given the noisy observation \mathbf{y}) is the conditional mean, which can be computed by [18]

$$E\{x_c|\mathbf{y}\} = \int_0^\infty E\{x_c|\mathbf{y}, z\}p(z|\mathbf{y})dz.$$
 (5)

The right hand side of Eq. (5) is a one-dimensional integral over z, where two components, $E\{x_c|\mathbf{y},z\}$ and $p(z|\mathbf{y})$, need to be computed for each given z. The first component $E\{x_c|\mathbf{y},z\}$ is linear based on the facts that \mathbf{w} is Gaussian and \mathbf{x} is also Gaussian when conditioned on z. In particular, we have

$$E\{\mathbf{x}|\mathbf{y},z\} = z\mathbf{C}_{\mathbf{u}}(z\mathbf{C}_{\mathbf{u}} + \mathbf{C}_{\mathbf{w}})^{-1}\mathbf{y},$$
 (6)

where $\mathbf{C_u}$ can be estimated from the observed noisy covariance matrix by $\mathbf{C_u} = \mathbf{C_y} - \mathbf{C_w}$. Equation (6) gives a full estimate of the original coefficient vector \mathbf{x} for any given z. For our purpose here, only the center coefficient x_c is of our interest, which leads to significant simplifications of the computation [18]. The second component in the integral in Eq. (5), i.e., the posterior density $p(z|\mathbf{y})$, can also be estimated by Bayes' rule: $p(z|\mathbf{y}) \propto p(\mathbf{y}|z)p_z(z)$, where $p_z(z)$ can be calculated using Eq. (3). Given Eq. (4) and the facts that both \mathbf{u} and \mathbf{w} are zero-mean Gaussian, it is straightforward to derive that $p(\mathbf{y}|z)$ can be computed as a Gaussian function of zero-mean and covariance $\mathbf{C_y} = z\mathbf{C_u} + \mathbf{C_w}$ [18].

III. SPATIOTEMPORAL GSM VIDEO DENOISING

Figure 1 illustrates the diagram of the proposed video denoising system. The denoising of the current frame involves not only the frame itself, but also a set of adjacent past and future frames. Motion estimation is performed between the current frame and the past/future frames. Details on the motion estimation algorithm will be given in Section IV. The estimation results are used for global motion compensation. Wavelet transform is then applied to the current frame as

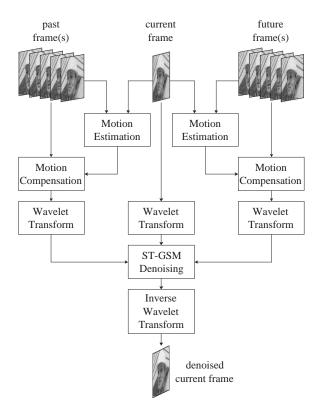


Fig. 1. Diagram of the proposed ST-GSM video denoising algorithm.

well as the motion compensated past and future frames. The wavelet transform is a linear multi-resolution analysis tool that decomposes an image signal into multiple subbands, each with a different characteristic scale and orientation. An excellent description can be found in [40]. Next, wavelet coefficient vectors are formed from a spatiotemporal neighborhood, and an ST-GSM denoising method similar to the BLS estimator discussed in Section II is employed. Finally, an inverse wavelet transform is applied to the denoised wavelet coefficients to create a denoised current frame.

At the core of the proposed video denoising system shown in Fig. 1 is the ST-GSM denoising algorithm, where the key is to include temporal neighborhoods in the wavelet coefficient vector. This is illustrated in Fig. 2, where a coefficient vector of length $N_1 \times N_2 \times N_f$ is formed by a set of spatial neighboring wavelet coefficients from the same scale and the same orientation but across a number of past and future frames. Here N_1 and N_2 denote the spatial dimensions and N_f is the total number of frames involved (including past, current and future frames). In the illustration given in Fig. 2, $N_1 = N_2 = N_f = 3$, but our experiments show that a larger number of N_f often leads to better denoising results, and we empirically set $N_f = 9$ in all of our denoising experiments.

There are a number of practical issues in our implementation of the proposed ST-GSM denoising algorithm:

- We choose an 8-orientation, 4-scale steerable pyramid [41] for wavelet decomposition. The steerable pyramid is an overcomplete wavelet transform that avoids aliasing in subbands.
- After global motion compensation, blank regions are

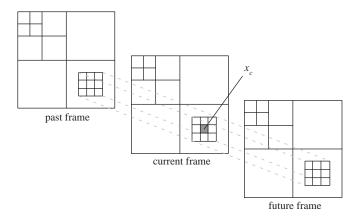


Fig. 2. The formation of $3\times3\times3$ wavelet coefficient neighborhood over three yideo frames

created at the image boundaries. We fill them by mirror replication of the shifted frame.

- The BLS denoising of coefficient x_c in Eq.(5) involves a 1D integration over z. We implement this (using default parameters as in [18]) by sampling z in logarithm domain by 13 points in the range of [-20.5, 3.5].
- For convenience, the first frames in a video sequence are denoised using the available past frames only. For example, when denoising the third frame, only 2 past frames and 4 future frames are involved. Similar strategy is employed in denoising the last frames.

IV. Noise-Robust Motion Estimation

One of the challenges in the implementation of the above algorithm is to estimate motion in the presence of noise. Here we propose a simple but reliable noise-robust cross correlation method for global motion estimation at integer pixel precision. The limitation of using global motion estimation is that it cannot account for rotation, zooming and local motion. However, local motion compensation processes (such as those based on block matching and optical flow) are spatially adaptive nonlinear operators, which significantly change noise statistics and affect the adequacy of the Gaussain noise model assumed in GSM denoising. As a result, the BLS-GSM denoising estimator becomes invalid. Therefore, we restrict our motion estimation to be global in our current implementation of ST-GSM.

Let $f_1(\mathbf{v})$ and $f_2(\mathbf{v})$ represent two image frames, where \mathbf{v} is a spatial integer index vector for the underlying 2D rectangular image lattice. A classical approach to estimating a global motion vector between the two frames is the cross correlation (CC) method [34–37], which is based on the observation that when $f_2(\mathbf{v})$ is a shifted version of $f_1(\mathbf{v})$, the position of the peak in the cross correlation function between $f_1(\mathbf{v})$ and $f_2(\mathbf{u})$ corresponds to the motion vector. Despite the simplicity of the idea, the computation of the cross correlation function is often costly. An equivalent but more efficient approach is to use the Fourier transform method: Let $F(\omega) = \mathcal{F}\{f(\mathbf{v})\}$ represents the 2D Fourier transform of an image frame, where \mathcal{F} denotes the Fourier transform operator. Then the cross

correlation function can be computed as:

$$k_{cc}(\mathbf{v}) = \mathscr{F}^{-1}\left\{Y(\boldsymbol{\omega})\right\}\,,\tag{7}$$

where $Y(\omega) = F_1(\omega)F_2^*(\omega)$. The estimated motion vector is given by

$$\mathbf{v}_{opt} = \operatorname*{argmax}_{\mathbf{v}} k_{cc}(\mathbf{v}) \,. \tag{8}$$

An interesting variation of this approach is the phase correlation (PC) method [42–44], where the Fourier spectrum is normalized in the frequency domain to have unit energy across all frequencies. The phase correlation function is given by

$$k_{pc}(\mathbf{v}) = \mathscr{F}^{-1} \left\{ \frac{Y(\boldsymbol{\omega})}{|Y(\boldsymbol{\omega})|} \right\}$$
 (9)

To have a close look, let us assume that $f_2(\mathbf{v})$ is simply a shifted version of $f_1(\mathbf{v})$, i.e., $f_2(\mathbf{v}) = f_1(\mathbf{v} - \Delta \mathbf{v})$. Based on the shifting property of the Fourier transform, we have $F_2(\boldsymbol{\omega}) = F_1(\boldsymbol{\omega}) \exp\{-j\boldsymbol{\omega}^T\Delta\mathbf{v}\}$ and $Y(\boldsymbol{\omega}) = |F_1(\boldsymbol{\omega})|^2 \exp\{j\boldsymbol{\omega}^T\Delta\mathbf{v}\}$, and thus

$$k_{pc}(\mathbf{v}) = \mathscr{F}^{-1}\left\{\exp\{j\boldsymbol{\omega}^T\Delta\mathbf{u}\}\right\} = \delta(\mathbf{v} + \Delta\mathbf{v}),$$
 (10)

which creates an impulse at the true motion vector position and is zero everywhere else.

Both CC and PC methods were designed with the assumption that there is no noise in the images. With the presence of noise, their performance degrades. Our noise-robust cross correlation (NRCC) function is defined as

$$k_{nrcc}(\mathbf{v}) = \mathscr{F}^{-1} \left\{ Y(\boldsymbol{\omega}) \left(1 - \frac{|N(\boldsymbol{\omega})|^2}{|Y(\boldsymbol{\omega})|} \right) \right\},$$
 (11)

where $|N(\omega)|^2$ is the noise power spectrum (in the case of white noise, $|N(\omega)|^2$ is a constant). To better understand this, it is useful to formulate the three approaches (PC, CC and NRCC) using a unified framework. In particular, each method can be viewed a specific weighting scheme in the Fourier domain:

$$k(\mathbf{v}) = \mathscr{F}^{-1} \left\{ W(\boldsymbol{\omega}) \exp\{j\boldsymbol{\omega}^T \Delta \mathbf{v}\} \right\}, \qquad (12)$$

where the differences lie in the definition of the weighting function $W(\omega)$:

$$W_{pc}(\omega) \equiv 1,$$

$$W_{cc}(\omega) = |F_1(\omega)|^2,$$

$$W_{nrcc}(\omega) = |F_1(\omega)|^2 - |N(\omega)|^2.$$
(13)

The PC method assigns uniform weights to all frequencies, the CC method assigns the weights based on the total signal power (which is the sum of signal and noise power), while the NRCC method assigns the weights proportional to the signal power only (by removing the noise power part). It converges to the CC method when the images are noise-free. We tested the PC, CC and NRCC methods at different noise levels and use Root Mean Squared (RMS) error between the true shift and estimated shift to evaluate their performance. Figure 3 shows our test results by estimating the shift between a one-dimensional signal (extracted from one row of the "Einstein" image) and a shifted version of it using the three methods. It appears that the CC and NRCC methods perform much better at all noise levels, and NRCC leads to the best performance.

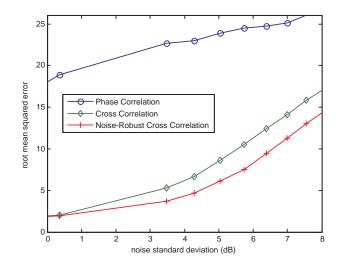


Fig. 3. RMS error comparison of phase correlation, cross correlation and noise-robust cross correlation based motion estimation algorithms as a function of noise level (normalized noise standard deviation in decibel).

V. VALIDATION

The video denoising algorithms were tested using publicly available video sequences [45] contaminated with additive white Gaussian noise. All video sequences are in YCrCb 4:2:0 format, but only the denoising results of the luma channel are reported here for algorithm validation. Two objective criteria, namely the peak signal-to-noise-ratio (PSNR) and the structural similarity index (SSIM) [46–48], were employed to provide quantitative quality evaluations of the denoising results. Specifically, PSNR is defined as

$$PSNR = 10\log_{10}\left(\frac{L^2}{MSF}\right),\tag{14}$$

where L is the dynamic range of the image (for 8bits/pixel images, L=255) and MSE is the mean squared error between the original and distorted images. SSIM is first calculated within local windows using

SSIM(
$$\mathbf{x}, \mathbf{y}$$
) = $\frac{(2\mu_x \mu_y + C_1)(2\sigma_{xy} + C_2)}{(\mu_x^2 + \mu_y^2 + C_1)(\sigma_x^2 + \sigma_y^2 + C_2)}$, (15)

where \mathbf{x} and \mathbf{y} are the image patches extracted from the local window from the original and distorted images, respectively. μ_x , σ_x^2 and σ_{xy} are the mean, variance, and cross-correlation computed within the local window, respectively. The overall SSIM score of a video frame is computed as the average local SSIM scores. PSNR is the mostly widely used quality measure in the literature, but has been criticized for not correlating well with human visual perception [49]. SSIM is believed to be a better indicator of perceived image quality [49]. It also supplies a quality map that indicates the variations of images quality over space. The final PSNR and SSIM results for a denoised video sequence are computed as the frame average of the full sequence, after clipping the denoised pixels to the range of [0, 255].

Three experiments were carried out to validate various aspects of the proposed ST-GSM algorithm. In the first experiment, we verify the usefulness of including temporal wavelet neighbors in forming the coefficient vector in GSM

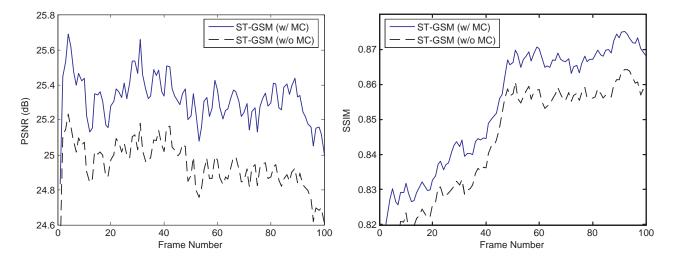


Fig. 4. PSNR and SSIM comparisons of ST-GSM denoising with and without motion estimation/motion compensation. (a) PSNR plot of denoised "Garden" sequence corrupted by noise with $\sigma=30$; (b) SSIM plot of denoised "Garden" sequence corrupted by noise with $\sigma=30$.

denoising. In particular, we compare the proposed algorithm against GSM denoising applied to each individual video frame independently (abbreviated as 2D-GSM [18]). The results are shown in Table I, where ST-GSM performs consistently better for all test sequences at all noise levels.

Second, the effectiveness of motion estimation/motion compensation in the proposed algorithm is tested. Figure 4 shows the results of a straightforward comparison, where the same ST-GSM algorithm was applied to denoise the same video sequence, but with and without the involvement of the motion estimation/motion compensation stages. The PSNR and SSIM results computed on a frame-by-frame basis show that motion estimation/compensation leads to markable improvement. Similar results were observed when denoising other video sequences, especially those video segments that involve significant amount of global motion.

In the third experiment, we compare ST-GSM against other video denoising algorithms, including both baseline and stateof-the-art schemes. These include 1) Wiener2D: MATLAB's Wiener2 function (a spatially adaptive Wiener filter) applied on a frame-by-frame basis; 2) Wiener3D: our implementation of a simple extension of Wiener2 to three dimensions, with a window size of $3 \times 3 \times 3$; 3) WRSTF [5]: wavelet-domain reliability-based spatio-temporal filtering; 4) SEQWT [10]: sequential wavelet domain temporal filtering; 5) 3DWTF [6]: 3D dual tree wavelet transform denoising; 6) IFSM [9]: interframe statistical modeling; 7) 3DSWDCT [7]: 3D sliding window Discrete Cosine Transform; 8) VBM3D [14]: block matching and 3D filtering; 9) 2D-GSM [18]: as in the first experiment described above; 10) STA [15]: space-time adaptive patch-based filtering; 11) K-SVD [16]: video denoising based on sparse and redundant representation; and 12) 3D-patch [17]: Bayesian variational 3D patch-based method. The results for WRSTF, SEQWT and 3DWTF were obtained from the processed video sequences available at [50]. The PSNR results of STA, K-SVD and 3D-patch denoising are directed cited from the corresponding publications [15–17]. We performed the denoising experiments for the rest of the methods.

Table I compares PSNR and SSIM results of 10 denoising algorithms for six video sequences at five noise levels. Table II shows PSNR and SSIM performance averaged over sequences. Table III presents PSNR comparisons with STA, K-SVD and 3D-patch methods. It can be observed that the proposed ST-GSM algorithm demonstrates competitive performance when compared with the state-of-the-art. Finally, Fig. 5 demonstrates the visual effects of different denoising algorithms. Specifically, we show a frame extracted from the "Miss America" sequence, together with a noisy version of the same frame, and the denoised frames obtained by four video denoising algorithms. It can be seen that ST-GSM is quite effective at suppressing background noise while maintaining the edge and texture details and thus the structural information of the objects in the scene. This is further verified by examining the SSIM quality maps of the corresponding frames.

VI. CONCLUSION

We propose a wavelet-domain ST-GSM model for natural video signals and applied it to the restoration of video signals corrupted by additive white Gaussian noise. We found that applying motion estimation/motion compensation is effective in enhancing the correlations between temporal neighboring wavelet coefficients and thus improving the performance of ST-GSM denoising. We propose a Fourier domain NRCC scheme to provide reliable motion estimation in the presence of noise. Our experimental comparisons with state-of-the-art algorithms show that ST-GSM is competitive in terms of both subjective and objective (PSNR and SSIM) evaluations.

There are a number of potential improvements and extensions that may be done in the future. First, the current implementation of ST-GSM denoising is rather slow. A rough estimate for the computational complexity can be inferred from the observation that our un-optimized Matlab code took 120 seconds per CIF frame on an Intel 2.4 GHz workstation. Both algorithmic and software optimizations are needed to

TABLE I
PSNR and SSIM [47] Comparisons of video denoising algorithms for 6 video sequences at 5 noise levels

video sequence	Foreman				Salesman				Miss America						
noise std (σ)	10	15	20	50	100	10	15	20	50	100	10	15	20	50	100
								Result							
Wiener2D	33.14	30.46	28.55	23.09	16.39	31.97	29.51	27.80	21.31	13.09	34.51	31.64	29.56	21.76	12.84
Wiener3D	29.54	29.26	28.87	25.35	15.28	29.59	29.30	28.88	22.92	13.50	36.95	35.60	34.06	23.39	13.27
WRSTF [5]		33.37	31.82	NA	NA	35.54		32.00	NA	NA	37.82		34.79	NA	NA
SEQWT [10]	NA	NA	NA	NA	NA		30.59		NA	NA	NA	NA	NA	NA	NA
3DWTF [6]	NA	NA	NA	NA	NA	34.96	33.33		NA	NA	NA	NA	NA	NA	NA
IFSM [9]			30.50			34.22		30.22				35.41			22.49
3DSWDCT [7]	36.17	34.46	33.07	27.33	21.46	36.98	35.12	33.75	28.82	22.30	38.87	37.72	36.74	32.88	23.43
VBM3D [14]	37.17	35.57	34.36	28.95	21.36	38.33		35.12			40.29	39.30	38.54	33.39	22.81
2D-GSM [18]	35.05	33.10	31.70	26.41	20.65	33.80	31.73	30.28	24.95	20.32	38.52	37.14	36.14	30.49	22.16
ST-GSM			33.72			38.04	36.03				40.58			31.62	
								M Res							
Wiener2D	0.860	0.774	0.694	0.425	0.164	0.859	0.778		0.340	0.074	0.818	0.704	0.602	0.214	0.044
Wiener3D	0.865	0.839	0.808	0.582	0.103	0.839	0.818	0.785	0.425	0.066	0.908	0.868	0.808	0.286	0.040
WRSTF [5]		0.877	0.841	NA	NA		0.901		NA	NA	0.908	0.877	0.846	NA	NA
SEQWT [10]	NA	NA	NA	NA	NA	0.900	0.846	0.796	NA	NA	NA	NA	NA	NA	NA
3DWTF [6]	NA	NA	NA	NA	NA	0.900		0.750	NA	NA	NA	NA	NA	NA	NA
IFSM [9]	0.886	0.836	0.793	0.667	0.666	0.904	0.851	0.801	0.609	0.488	0.904	0.857	0.812	0.780	0.804
3DSWDCT [7]		0.907	0.753	0.769	0.708	0.955	0.930		0.772	0.400	0.946	0.928	0.909	0.750	0.829
VBM3D [14]	0.935	0.917	0.903	0.837	0.699	0.960	0.945	0.925	0.771	0.538	0.947	0.939	0.933	0.905	0.789
2D-GSM [18]	0.500	0.889			0.672	0.909	0.865	****	0.611	0.336	0.936	0.922		0.874	
ST-GSM	0.937	0.917		0.820	0.715	0.960			0.727	0.496	0.952	0.943		0.892	0.823
DI GBIT	0.707	0.717	0.701	0.020	0.710	0.200	0.711	0.723	0.727	0.170	0.70=	0.7 .0	0.700	0.072	0.023
video sequence			Tennis					Garden					Footbal	1	
video sequence noise std (σ)	10	15	Tennis 20	50	100	10	15			100	10	15	Footbal 20		100
video sequence noise std (σ)	10	15	Tennis 20	50	100	10	15	Garden 20 Result	50	100	10		Footbal 20	50	100
noise std (σ)			20			I	15 PSNR	20 Result	50 s (dB)			15	20	50	
noise std (σ) Wiener2D	31.07	28.55	26.78	20.65	15.46	29.73	15 PSNR 26.77	20 Result 24.80	50 s (dB) 18.09	14.17	25.07	15 24.76	24.37	21.19	15.23
noise std (σ) Wiener2D Wiener3D	31.07 22.88	28.55 22.81	20 26.78 22.71	20.65 21.50	15.46 15.14	29.73 18.36	15 PSNR 26.77 18.33	20 Result 24.80 18.30	50 s (dB) 18.09 17.82	14.17 13.49	25.07 21.97	24.76 21.91	20 24.37 21.84	50 21.19 20.82	15.23 14.73
noise std (σ) Wiener2D Wiener3D WRSTF [5]	31.07 22.88 33.68	28.55 22.81 31.35	20 26.78 22.71 29.71	20.65 21.50 NA	15.46 15.14 NA	29.73 18.36 30.59	15 PSNR 26.77 18.33 27.95	20 Result 24.80 18.30 26.13	50 s (dB) 18.09 17.82 NA	14.17 13.49 NA	25.07 21.97 NA	24.76 21.91 NA	24.37 21.84 NA	21.19 20.82 NA	15.23 14.73 NA
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10]	31.07 22.88 33.68 31.19	28.55 22.81 31.35 29.14	26.78 22.71 29.71 27.59	20.65 21.50 NA NA	15.46 15.14 NA NA	29.73 18.36 30.59 29.30	15 PSNR 26.77 18.33 27.95 26.43	20 Result 24.80 18.30 26.13 24.38	50 s (dB) 18.09 17.82 NA NA	14.17 13.49 NA NA	25.07 21.97	24.76 21.91 NA NA	24.37 21.84 NA NA	21.19 20.82 NA NA	15.23 14.73 NA NA
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6]	31.07 22.88 33.68 31.19 31.96	28.55 22.81 31.35 29.14 29.91	26.78 22.71 29.71 27.59 28.56	20.65 21.50 NA NA NA	15.46 15.14 NA NA NA	29.73 18.36 30.59 29.30 30.25	15 PSNR 26.77 18.33 27.95 26.43 27.70	20 Result 24.80 18.30 26.13 24.38 25.95	50 s (dB) 18.09 17.82 NA NA NA	14.17 13.49 NA NA NA	25.07 21.97 NA NA NA	24.76 21.91 NA NA NA	24.37 21.84 NA NA NA	21.19 20.82 NA NA NA	15.23 14.73 NA NA NA
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9]	31.07 22.88 33.68 31.19 31.96 32.41	28.55 22.81 31.35 29.14 29.91 30.10	26.78 22.71 29.71 27.59 28.56 28.56	20.65 21.50 NA NA NA NA 23.81	15.46 15.14 NA NA NA NA 20.76	29.73 18.36 30.59 29.30 30.25 30.05	15 PSNR 26.77 18.33 27.95 26.43 27.70 27.25	20 Result 24.80 18.30 26.13 24.38 25.95 25.40	50 s (dB) 18.09 17.82 NA NA NA NA 20.23	14.17 13.49 NA NA NA NA 16.50	25.07 21.97 NA NA NA NA 31.23	24.76 21.91 NA NA NA NA 28.78	24.37 21.84 NA NA NA NA 27.15	21.19 20.82 NA NA NA 22.62	15.23 14.73 NA NA NA NA 19.79
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7]	31.07 22.88 33.68 31.19 31.96 32.41 33.83	28.55 22.81 31.35 29.14 29.91 30.10 31.79	26.78 22.71 29.71 27.59 28.56 28.56 30.50	20.65 21.50 NA NA NA 23.81 26.16	15.46 15.14 NA NA NA 20.76 22.00	29.73 18.36 30.59 29.30 30.25 30.05 31.80	15 PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70	50 s (dB) 18.09 17.82 NA NA NA 20.23 21.50	14.17 13.49 NA NA NA 16.50 16.95	25.07 21.97 NA NA NA 31.23 32.32	24.76 21.91 NA NA NA	24.37 21.84 NA NA NA 27.15 28.47	21.19 20.82 NA NA NA 22.62 23.78	15.23 14.73 NA NA NA
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88	26.78 22.71 29.71 27.59 28.56 28.56 30.50 31.49	20.65 21.50 NA NA NA 23.81 26.16 27.21	15.46 15.14 NA NA NA 20.76 22.00 21.06	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74	50 s (dB) 18.09 17.82 NA NA NA 20.23 21.50 22.52	14.17 13.49 NA NA NA 16.50 16.95 16.51	25.07 21.97 NA NA NA 31.23 32.32 33.09	24.76 21.91 NA NA NA 28.78 29.97 30.90	24.37 21.84 NA NA NA 27.15 28.47 29.36	21.19 20.82 NA NA NA 22.62 23.78 24.58	15.23 14.73 NA NA NA 19.79 19.99 20.04
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82	28.55 22.81 31.35 29.14 29.91 30.10 31.79	26.78 22.71 29.71 27.59 28.56 28.56 30.50 31.49	20.65 21.50 NA NA NA 23.81 26.16	15.46 15.14 NA NA NA 20.76 22.00	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40	15 PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76	50 s (dB) 18.09 17.82 NA NA NA 20.23 21.50 22.52 19.98	14.17 13.49 NA NA NA 16.50 16.95	25.07 21.97 NA NA NA 31.23 32.32	24.76 21.91 NA NA NA 28.78 29.97 30.90	24.37 21.84 NA NA NA 27.15 28.47 29.36 27.74	21.19 20.82 NA NA NA 22.62 23.78 24.58	15.23 14.73 NA NA NA 19.79 19.99
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87	20 26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65	20.65 21.50 NA NA NA 23.81 26.16 27.21 24.36	15.46 15.14 NA NA NA 20.76 22.00 21.06 20.10	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23	14.17 13.49 NA NA NA 16.50 16.95 16.51 16.04	25.07 21.97 NA NA 31.23 32.32 33.09 31.33	24.76 21.91 NA NA NA 28.78 29.97 30.90 29.14	24.37 21.84 NA NA NA 27.15 28.47 29.36 27.74	21.19 20.82 NA NA NA 22.62 23.78 24.58 23.30	15.23 14.73 NA NA NA 19.79 19.99 20.04 19.31
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87	20 26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65	20.65 21.50 NA NA NA 23.81 26.16 27.21 24.36	15.46 15.14 NA NA NA 20.76 22.00 21.06 20.10	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78	25.07 21.97 NA NA 31.23 32.32 33.09 31.33	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87	24.37 21.84 NA NA 27.15 28.47 29.36 27.74 28.36	21.19 20.82 NA NA NA 22.62 23.78 24.58 23.30	15.23 14.73 NA NA NA 19.79 19.99 20.04 19.31
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97	26.78 22.71 29.71 27.59 28.56 28.56 30.50 31.49 28.65 30.59	20.65 21.50 NA NA 23.81 26.16 27.21 24.36 25.85	15.46 15.14 NA NA 20.76 22.00 21.06 20.10 20.40	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults	14.17 13.49 NA NA NA 16.50 16.95 16.51 16.04	25.07 21.97 NA NA 31.23 32.32 33.09 31.33 32.11	24.76 21.91 NA NA NA 28.78 29.97 30.90 29.14	24.37 21.84 NA NA NA 27.15 28.47 29.36 27.74	21.19 20.82 NA NA 22.62 23.78 24.58 23.30 23.58	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM Wiener2D Wiener3D	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97	26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65 30.59 0.625 0.539	20.65 21.50 NA NA 23.81 26.16 27.21 24.36 25.85	15.46 15.14 NA NA NA 20.76 22.00 21.06 20.10 20.40 0.092 0.045	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08 SS 0.853	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49 M Res 0.792 0.488	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults 0.413	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78	25.07 21.97 NA NA NA 31.23 32.32 33.09 31.33 32.11	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87 0.676 0.581	24.37 21.84 NA NA NA 27.15 28.47 29.36 27.74 28.36 0.643 0.568	21.19 20.82 NA NA NA 22.62 23.78 24.58 23.30 23.58	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM Wiener2D Wiener3D WRSTF [5]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05 0.813 0.577 0.897	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97 0.712 0.560	26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65 30.59 0.625 0.539 0.790	20.65 21.50 NA NA NA 23.81 26.16 27.21 24.36 25.85 0.288 0.363	15.46 15.14 NA NA NA 20.76 22.00 21.06 20.10 20.40	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08 SS 0.853 0.500	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49 M Res 0.792 0.488 0.889	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults 0.413 0.391	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78	25.07 21.97 NA NA 31.23 32.32 33.09 31.33 32.11	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87	24.37 21.84 NA NA NA 27.15 28.47 29.36 27.74 28.36	21.19 20.82 NA NA NA 22.62 23.78 24.58 23.30 23.58	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM Wiener2D Wiener3D WRSTF [5] SEQWT [10]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05 0.813 0.577 0.897	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97 0.712 0.560 0.839	26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65 30.59 0.625 0.539 0.790 0.716	20.65 21.50 NA NA NA 23.81 26.16 27.21 24.36 25.85 0.288 0.363 NA	15.46 15.14 NA NA 20.76 22.00 21.06 20.10 20.40 0.092 0.045 NA	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08 SS 0.853 0.500 0.922	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49 M Res 0.792 0.488 0.889 0.842	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults 0.413 0.391 NA	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78	25.07 21.97 NA NA 31.23 32.32 33.09 31.33 32.11 0.703 0.590 NA	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87 0.676 0.581 NA	24.37 21.84 NA NA NA 27.15 28.47 29.36 27.74 28.36 0.643 0.568 NA	21.19 20.82 NA NA NA 22.62 23.78 24.58 23.30 23.58 0.434 0.437 NA	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66 0.157 0.071 NA
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05 0.813 0.577 0.897 0.842	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97 0.712 0.560 0.839 0.772	26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65 30.59 0.625 0.539 0.790	20.65 21.50 NA NA NA 23.81 26.16 27.21 24.36 25.85 0.288 0.363 NA NA	15.46 15.14 NA NA 20.76 22.00 21.06 20.10 20.40 0.092 0.045 NA NA	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48 0.916 0.510 0.953 0.941	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08 SS 0.853 0.500 0.922 0.893 0.872	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49 M Res 0.792 0.488 0.889 0.842	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults 0.413 0.391 NA NA	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78 0.165 0.041 NA NA	25.07 21.97 NA NA 31.23 32.32 33.09 31.33 32.11 0.703 0.590 NA NA	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87 0.676 0.581 NA NA	24.37 21.84 NA NA 27.15 28.47 29.36 27.74 28.36 0.643 0.568 NA NA	21.19 20.82 NA NA 22.62 23.78 24.58 23.30 23.58 0.434 0.437 NA NA	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66 0.157 0.071 NA NA
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05 0.813 0.577 0.897 0.842 0.856 0.855	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97 0.712 0.560 0.839 0.772 0.793 0.776	26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65 30.59 0.625 0.539 0.790 0.716 0.740 0.709	20.65 21.50 NA NA NA 23.81 26.16 27.21 24.36 25.85 0.288 0.363 NA NA NA O.485	15.46 15.14 NA NA 20.76 22.00 21.06 20.10 20.40 0.092 0.045 NA NA NA O.458	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48 0.916 0.953 0.941 0.909 0.927	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08 SS 0.853 0.500 0.922 0.893 0.872 0.882	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49 M Res 0.792 0.488 0.889 0.842 0.840 0.837	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults 0.413 0.391 NA NA NA 0.623	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78 0.165 0.041 NA NA NA O.344	25.07 21.97 NA NA 31.23 32.32 33.09 31.33 32.11 0.703 0.590 NA NA NA 0.884	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87 0.676 0.581 NA NA NA O.813	24.37 21.84 NA NA 27.15 28.47 29.36 27.74 28.36 0.643 0.568 NA NA NA O.749	21.19 20.82 NA NA 22.62 23.78 24.58 23.30 23.58 0.434 0.437 NA NA NA O.510	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66 0.157 0.071 NA NA NA O.370
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05 0.813 0.577 0.897 0.842 0.856 0.855 0.894	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97 0.712 0.560 0.839 0.772 0.793	26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65 30.59 0.625 0.539 0.790 0.716 0.740	20.65 21.50 NA NA NA 23.81 26.16 27.21 24.36 25.85 0.288 0.363 NA NA	15.46 15.14 NA NA 20.76 22.00 21.06 20.10 20.40 0.092 0.045 NA NA	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48 0.916 0.510 0.953 0.941 0.909	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08 SS 0.853 0.500 0.922 0.893 0.872	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49 M Res 0.792 0.488 0.889 0.842 0.840	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults 0.413 0.391 NA NA	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78 0.165 0.041 NA NA	25.07 21.97 NA NA 31.23 32.32 33.09 31.33 32.11 0.703 0.590 NA NA	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87 0.676 0.581 NA NA	24.37 21.84 NA NA 27.15 28.47 29.36 27.74 28.36 0.643 0.568 NA NA	21.19 20.82 NA NA 22.62 23.78 24.58 23.30 23.58 0.434 0.437 NA NA	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66 0.157 0.071 NA NA
noise std (σ) Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05 0.813 0.577 0.897 0.842 0.856 0.855 0.894	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97 0.712 0.560 0.839 0.772 0.793 0.776 0.834	26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65 30.59 0.625 0.539 0.790 0.716 0.740 0.709 0.790	20.65 21.50 NA NA 23.81 26.16 27.21 24.36 25.85 0.288 0.363 NA NA NA 0.485 0.620	15.46 15.14 NA NA 20.76 22.00 21.06 20.10 20.40 0.092 0.045 NA NA NA 0.458 0.513	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48 0.916 0.953 0.941 0.909 0.927 0.959	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08 SS 0.853 0.500 0.922 0.893 0.872 0.882 0.931	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49 M Res 0.792 0.488 0.889 0.842 0.840 0.837 0.900	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults 0.413 0.391 NA NA NA 0.623 0.688	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78 0.165 0.041 NA NA NA 0.344 0.396	25.07 21.97 NA NA 31.23 32.32 33.09 31.33 32.11 0.703 0.590 NA NA 0.884 0.911	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87 0.676 0.581 NA NA O.813 0.851	24.37 21.84 NA NA 27.15 28.47 29.36 27.74 28.36 0.643 0.568 NA NA NA 0.749 0.801	21.19 20.82 NA NA 22.62 23.78 24.58 23.30 23.58 0.434 0.437 NA NA NA 0.510 0.595	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66 0.157 0.071 NA NA 0.370 0.398
wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7] VBM3D [14] 2D-GSM [18] ST-GSM Wiener2D Wiener3D WRSTF [5] SEQWT [10] 3DWTF [6] IFSM [9] 3DSWDCT [7]	31.07 22.88 33.68 31.19 31.96 32.41 33.83 34.89 31.82 34.05 0.813 0.577 0.897 0.842 0.856 0.855 0.894 0.901 0.831	28.55 22.81 31.35 29.14 29.91 30.10 31.79 32.88 29.87 31.97 0.712 0.560 0.839 0.772 0.793 0.776 0.834 0.847 0.758	26.78 22.71 29.71 27.59 28.56 30.50 31.49 28.65 30.59 0.625 0.539 0.790 0.716 0.740 0.709 0.800	20.65 21.50 NA NA NA 23.81 26.16 27.21 24.36 25.85 0.288 0.363 NA NA NA 0.485 0.620 0.640 0.577	15.46 15.14 NA NA 20.76 22.00 21.06 20.10 20.40 0.092 0.045 NA NA 0.458 0.513 0.478 0.456	29.73 18.36 30.59 29.30 30.25 30.05 31.80 32.54 30.40 31.48 0.916 0.510 0.953 0.941 0.909 0.927 0.959 0.962 0.939	PSNR 26.77 18.33 27.95 26.43 27.70 27.25 29.40 30.30 27.65 29.08 SS 0.853 0.500 0.922 0.893 0.872 0.882 0.931 0.940 0.899	20 Result 24.80 18.30 26.13 24.38 25.95 25.40 27.70 28.74 25.76 27.49 M Res 0.792 0.488 0.889 0.842 0.840 0.837 0.900 0.916	50 s (dB) 18.09 17.82 NA NA 20.23 21.50 22.52 19.98 22.23 ults 0.413 0.391 NA NA 0.623 0.688 0.738 0.611	14.17 13.49 NA NA 16.50 16.95 16.51 16.04 16.78 0.165 0.041 NA NA 0.344 0.396	25.07 21.97 NA NA 31.23 32.32 33.09 31.33 32.11 0.703 0.590 NA NA 0.884 0.911 0.923	24.76 21.91 NA NA 28.78 29.97 30.90 29.14 29.87 0.676 0.581 NA NA 0.813 0.851 0.874 0.798	24.37 21.84 NA NA 27.15 28.47 29.36 27.74 28.36 0.643 0.568 NA NA 0.749 0.801 0.822	21.19 20.82 NA NA 22.62 23.78 24.58 23.30 23.58 0.434 0.437 NA NA 0.510 0.595 0.601 0.552	15.23 14.73 NA NA 19.79 19.99 20.04 19.31 19.66 0.157 0.071 NA NA 0.370 0.398 0.398

accelerate the algorithm. Second, the current algorithm is applied to each color channel independently. Denoising all color channels jointly by including color wavelet coefficient neighbors in GSM modeling may further improve the algorithm. Finally, there are advanced models in describing the statistical motion properties of natural video signals (e.g., [24]) that may be employed to impose stronger statistical prior in a Bayesian framework and in turn improve the denoising performance.

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TABLE III
PSNR COMPARISONS WITH LATEST VIDEO DENOISING ALGORITHMS

Sequence	Size	Input PSNR	STA [15]	K-SVD [16]	3D-Patch [17]	ST-GSM
Salesman	$176 \times 144 \times 449$	28	35.13	37.91	39.26	37.93
		24	32.60	35.59	36.35	35.17
Miss America	$176 \times 144 \times 108$	28	39.39	40.49	42.23	41.43
Suzie	$176 \times 144 \times 150$	28	37.07	37.96	38.40	38.36
		24	35.11	35.95	36.32	36.21
Trevor	$176 \times 144 \times 90$	28	36.68	38.10	38.31	37.17
		24	34.79	35.97	35.81	34.68
Foreman	$176 \times 144 \times 300$	28	34.94	37.86	36.88	36.85
		24	32.90	35.86	34.55	34.37

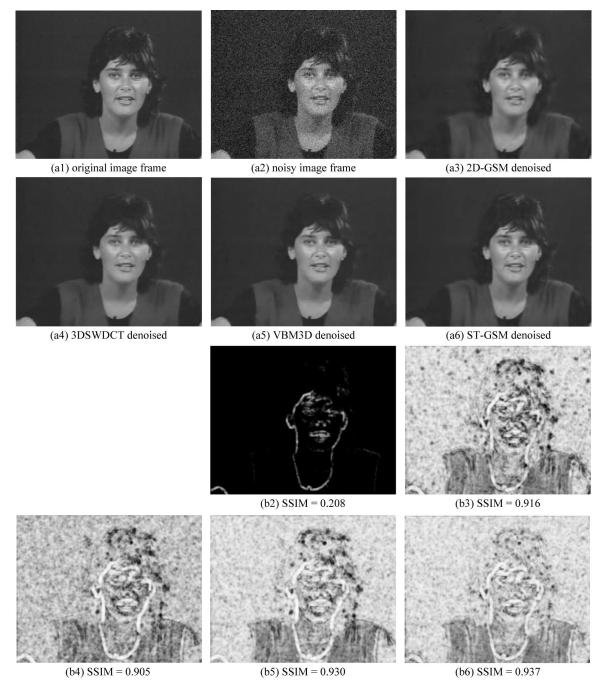


Fig. 5. Denoising results of Frame 80 in "Miss America" sequence corrupted with noise standard deviation $\sigma = 20$. (a1)-(a6): image frames in the original, noisy, and 2D-GSM [18], IFSM [9], VBM3D [14] and ST-GSM denoised sequences. (b2)-(b6): corresponding SSIM quality maps (brighter indicates larger SSIM value) with mean SSIM values.

TABLE II
AVERAGE PSNR AND SSIM [47] PERFORMANCE OF VIDEO DENOISING
ALGORITHMS AT 5 NOISE LEVELS

noise std (σ)	10	15	20	50	100		
	PSNR Results (dB)						
Wiener2D	30.92	28.62	26.98	21.02	14.53		
Wiener3D	26.55	26.20	25.78	21.97	14.24		
IFSM [9]	33.26	30.90	29.28	24.60	20.22		
3DSWDCT [7]	35.00	33.08	31.71	26.75	21.02		
VBM3D [14]	36.05	34.26	32.94	27.52	20.53		
2D-GSM [18]	33.49	31.44	30.05	24.92	19.76		
ST-GSM	35.50	33.56	32.21	26.33	20.30		
		SS	M Res	ults			
Wiener2D	0.828	0.750	0.677	0.352	0.116		
Wiener3D	0.715	0.694	0.666	0.414	0.061		
IFSM [9]	0.893	0.836	0.784	0.612	0.522		
3DSWDCT [7]	0.933	0.897	0.865	0.716	0.576		
VBM3D [14]	0.938	0.910	0.883	0.749	0.540		
2D-GSM [18]	0.900	0.855	0.820	0.668	0.507		
ST-GSM	0.934	0.905	0.880	0.738	0.538		

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